

11

5.9

Please refer to the picture of the heavy top on P. 209 or the picture on next page.

1

We are asked to derive the two constants of motion for the heavy top given by Eqs. 5.53 & 5.54 from the Euler's Equations.

$$5.53 \rightarrow I_3 (\dot{\psi} + \dot{\phi} \cos \theta) = \text{const}$$

$$5.54 \rightarrow (I_1 \sin^2 \theta + I_3 \cos^2 \theta) \dot{\phi} + I_3 \dot{\psi} \cos \theta = \text{const}$$

The Euler's Equations are given by:

$$(*) \begin{cases} I_1 \dot{\omega}_1 - (I_2 - I_3) \omega_2 \omega_3 = N_1 \\ I_2 \dot{\omega}_2 - (I_3 - I_1) \omega_3 \omega_1 = N_2 \\ I_3 \dot{\omega}_3 - (I_1 - I_2) \omega_1 \omega_2 = N_3 \end{cases}$$

- With the body axes (x_1, x_2, x_3) chosen as shown along the principle axes of the heavy top, we have:

$$\underline{I_1 = I_2 \neq I_3}$$

- By RHR, the torque \vec{N} can be shown to be along the line of nodes and

$$\underline{\begin{cases} N_1 = N \cos \psi \\ N_2 = -N \sin \psi \\ N_3 = 0 \end{cases}}$$

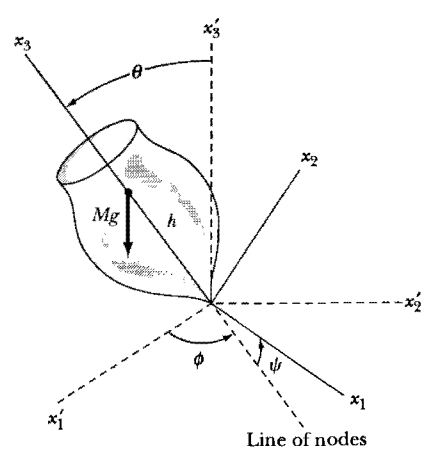


Figure: 11-15

Size: 13p9 x 14p6



Wadsworth
Thornton/Marion
Classical Dynamics 5/e
7973

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Note: one can also get the components of the force (by gravity) along body axes by multiplying the full Euler angle transformation matrix to $\begin{pmatrix} 0 \\ 0 \\ -mg \end{pmatrix}$.

(3)

$$\begin{aligned} \vec{F}_g &= \begin{pmatrix} \cos\psi & \sin\psi & 0 \\ -\sin\psi & \cos\psi & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta & \sin\theta \\ 0 & -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \cos\phi & \sin\phi & 0 \\ -\sin\phi & \cos\phi & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 0 \\ 0 \\ -mg \end{pmatrix} \\ &= \begin{pmatrix} \cos\psi & \sin\psi & 0 \\ -\sin\psi & \cos\psi & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 0 \\ -mg\sin\theta \\ -mg\cos\theta \end{pmatrix} = \begin{pmatrix} -mg\sin\theta\sin\psi \\ -mg\sin\theta\cos\psi \\ -mg\cos\theta \end{pmatrix} \end{aligned}$$

$$\begin{aligned} \text{Then, } \vec{N} &= \vec{r} \times \vec{F}_g \\ &= \begin{pmatrix} \hat{r}_1 & \hat{r}_2 & \hat{r}_3 \\ 0 & 0 & r \\ F_g^1 & F_g^2 & F_g^3 \end{pmatrix} = \begin{pmatrix} -rF_g^2 \\ rF_g^1 \\ 0 \end{pmatrix} \\ \vec{N} &= \begin{pmatrix} mg r \sin\theta \cos\psi \\ -mg r \sin\theta \sin\psi \\ 0 \end{pmatrix} = \begin{pmatrix} N \cos\psi \\ -N \sin\psi \\ 0 \end{pmatrix} \end{aligned}$$

where, $N = rmg\sin\theta$

Now, we can simplify the Euler's Equations $\textcircled{4}$:

$$\begin{cases} I_1 \dot{\omega}_1 - (I_1 - I_3) \omega_2 \omega_3 = N \cos \varphi \\ I_1 \dot{\omega}_2 - (I_3 - I_1) \omega_3 \omega_1 = -N \sin \varphi \\ I_3 \dot{\omega}_3 = 0 \end{cases}$$

- We can now use 4.87 to express the angular velocity $\vec{\omega}$ in the body axes in terms of the Euler's angle:

$$\begin{cases} \omega_1 = \dot{\phi} \sin \theta \sin \varphi + \dot{\theta} \cos \varphi \\ \omega_2 = \dot{\phi} \sin \theta \cos \varphi - \dot{\theta} \sin \varphi \\ \omega_3 = \dot{\phi} \cos \theta + \dot{\varphi} \end{cases}$$

- Substituting ω_3 into the 3rd equation, we immediately get

$$I_3 \frac{d}{dt} (\dot{\phi} \cos \theta + \dot{\varphi}) = 0$$

$$\frac{d}{dt} (I_3 (\dot{\phi} \cos \theta + \dot{\varphi})) = 0$$

$$\Rightarrow \underline{I_3 (\dot{\phi} \cos \theta + \dot{\varphi}) = \text{const}}$$

← one of the constant of motion
Eq. 5.53.

Now, substitute ω_1 & ω_2 into the first two equations:

(5)

$$\textcircled{1} \quad I_1 \frac{d}{dt} (\dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos \psi) - (I_1 - I_2) (\dot{\phi} \sin \theta \cos \psi - \dot{\theta} \sin \psi) \cdot (\dot{\phi} \cos \theta + \dot{\psi}) = N \cos \psi$$

$$\textcircled{2} \quad I_1 \frac{d}{dt} (\dot{\phi} \sin \theta \cos \psi - \dot{\theta} \sin \psi) - (I_3 - I_1) (\dot{\phi} \cos \theta + \dot{\psi}) \cdot (\dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos \psi) = -N \sin \psi$$

* The goal is to combine these two equations in such a way that the LHS will be a full time derivative and the RHS = 0. Then the expression inside the full time derivative will be a constant of motion.

$\textcircled{1} \sin \psi + \textcircled{2} \cos \psi :$

$$I_1 \left[\sin \psi \frac{d}{dt} (\dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos \psi) + \cos \psi \frac{d}{dt} (\dot{\phi} \sin \theta \cos \psi - \dot{\theta} \sin \psi) \right] - (I_1 - I_3) \left[(\sin \psi \dot{\phi} \sin \theta \cos \psi - \dot{\theta} \sin^2 \psi) \cdot (\dot{\phi} \cos \theta + \dot{\psi}) - \cos \psi (\dot{\phi} \cos \theta + \dot{\psi}) (\dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos^2 \psi) \right] = 0$$

↓ the $(I_1 - I_3)$ term

$$- (I_1 - I_3) \left[(\dot{\phi} \sin \theta \sin \psi \cos \psi) \cdot (\dot{\phi} \cos \theta + \dot{\psi}) - (\dot{\phi} \sin \theta \sin \psi \cos \psi) \cdot (\dot{\phi} \cos \theta + \dot{\psi}) - \dot{\theta} (\dot{\phi} \cos \theta + \dot{\psi}) (\sin^2 \psi + \cos^2 \psi) \right]$$

$$= (I_1 - I_3) \dot{\theta} (\dot{\phi} \cos \theta + \dot{\psi})$$

The I_1 (first) term =

$$\begin{aligned}
& I_1 \left[\sin\theta \left[(\dot{\phi} \sin\theta) \cancel{\cos\theta} \dot{\psi} + \frac{d}{dt}(\dot{\phi} \sin\theta) \sin\theta + \frac{d}{dt}(\dot{\theta} \cos\theta) \right] \right. \\
& \quad \left. + \cos\theta \left[-(\dot{\phi} \sin\theta) \cancel{\sin\theta} \dot{\psi} + \frac{d}{dt}(\dot{\phi} \sin\theta) \cos\theta - \frac{d}{dt}(\dot{\theta} \sin\theta) \right] \right] \\
& = I_1 \left[\frac{d}{dt}(\dot{\phi} \sin\theta) (\sin^2\theta + \cos^2\theta) + \sin\theta (-\dot{\theta} \sin\theta \dot{\psi} + \dot{\theta} \cos\theta) \right. \\
& \quad \left. - \cos\theta (\dot{\theta} \cos\theta \dot{\psi} + \dot{\theta} \sin\theta) \right] \\
& = I_1 \left[\frac{d}{dt}(\dot{\phi} \sin\theta) - \dot{\theta} \dot{\psi} \right]
\end{aligned}$$

- Putting these two terms back together, we have

$$I_1 \left[\frac{d}{dt}(\dot{\phi} \sin\theta) - \dot{\theta} \dot{\psi} \right] + (I_1 - I_3) \dot{\theta} (\dot{\phi} \cos\theta + \dot{\psi}) = 0$$

$$I_1 \frac{d}{dt}(\dot{\phi} \sin\theta) - I_1 \dot{\theta} \dot{\psi} + I_1 \dot{\theta} \dot{\phi} \cos\theta + I_1 \dot{\theta} \dot{\psi} - I_3 \dot{\theta} \dot{\phi} \cos\theta - I_3 \dot{\theta} \dot{\psi} = 0$$

$$I_1 \frac{d}{dt}(\dot{\phi} \sin\theta) + I_1 \dot{\theta} \dot{\phi} \cos\theta - I_3 \dot{\theta} (\dot{\phi} \cos\theta + \dot{\psi}) = 0$$

- Multiply $\sin\theta$ on both sides:

$$I_1 \sin\theta \frac{d}{dt}(\dot{\phi} \sin\theta) + I_1 \dot{\theta} \dot{\phi} \cos\theta \sin\theta - I_3 (\dot{\theta} \sin\theta) (\dot{\phi} \cos\theta + \dot{\psi}) = 0$$

$$\begin{aligned}
& \underbrace{I_1 \frac{d}{dt}(\sin\theta \dot{\phi} \sin\theta)}_{\downarrow} + I_3 \frac{d}{dt}(\cos\theta) (\dot{\phi} \cos\theta + \dot{\psi})
\end{aligned}$$

- We already know that $I_3 (\dot{\phi} \cos \theta + \dot{\psi}) = \text{const}$,

so that we can absorb it into the $\frac{d}{dt}(\cos \theta)$ term:

So, we have

$$I_1 \frac{d}{dt} (\sin \theta (\dot{\phi} \sin \theta)) + I_3 \frac{d}{dt} (\cos \theta (\dot{\phi} \cos \theta + \dot{\psi})) = 0$$

$$\Rightarrow \frac{d}{dt} [I_1 \dot{\phi} \sin^2 \theta + I_3 \dot{\phi} \cos^2 \theta + I_3 \cos \theta \dot{\psi}] = 0$$

Finally, we have

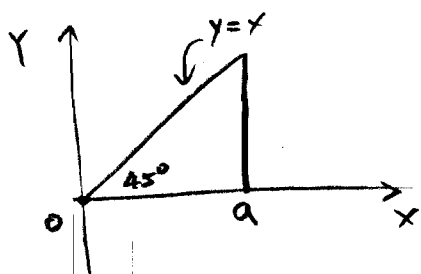
$$(I_1 \sin^2 \theta + I_3 \cos^2 \theta) \dot{\phi} + I_3 \dot{\psi} \cos \theta = \text{const}$$

which is the desired 2nd constant of motion.

5.15

For simplicity in integration, we will pick the following coordinate system for the rigid flat triangle.

①



← In this coordinate system, 2D integrals can be written simply as:

$$\int_0^a dx \int_0^x dy$$

- We will then use the parallel axis theorem to find the moments of inertia at the center of mass.

- Let the uniform mass density be σ per area, then the total mass M of the triangle is:

$$M = \sigma \int_0^a dx \int_0^x dy = \sigma \int_0^a x dx = \underline{\frac{1}{2} \sigma a^2}$$

- Now, we will find the CM:

$$\begin{aligned} Mx_{cm} &= \sigma \int_0^a \int_0^x x dx dy = \sigma \int_0^a x^2 dx \\ &= \frac{1}{3} \sigma a^3 = \underline{\frac{2}{3} aM} \end{aligned}$$

$$\begin{aligned} My_{cm} &= \sigma \int_0^a \int_0^x y dx dy = \sigma \int_0^a \frac{x^2}{2} dx \\ &= \frac{1}{6} \sigma a^3 = \underline{\frac{1}{3} aM} \end{aligned}$$

Now, we calculate the moments of inertia with respect to this coordinate system with the axis of rotation at the tip of the triangle (the origin). (2)

$$I_{xx} = \sigma \int_0^a \int_0^x y^2 dx dy = \sigma \int_0^a \frac{x^3}{3} dx = \frac{1}{12} \sigma a^4 = \frac{1}{6} M a^2$$

$$I_{yy} = \sigma \int_0^a \int_0^x x^2 dx dy = \sigma \int_0^a x^3 dx = \frac{1}{4} \sigma a^4 = \frac{1}{2} M a^2$$

$$I_{xy} = -\sigma \int_0^a \int_0^x xy dx dy = -\sigma \int_0^a x \frac{x^2}{2} dx = -\frac{1}{8} \sigma a^4 = -\frac{1}{4} M a^2$$

$$I_{zz} = \sigma \int_0^a \int_0^x (x^2 + y^2) dx dy = \sigma \int_0^a \left(x^2 \cdot x + \frac{x^3}{2} \right) dx$$

$$= \frac{1}{4} \sigma \frac{4}{3} a^4 = \frac{2}{3} M a^2$$

$$I_{xz} = \sigma \int_0^a \int_0^x x z_0 dx dy = \sigma \int_0^a \int_0^x y z_0 dx dy = 0$$

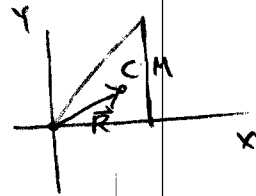
triangle lays flat at $z_0 = 0$.

$$\Rightarrow \underline{\underline{I}} \approx \begin{pmatrix} 1/6 & -1/4 & 0 \\ -1/4 & 1/2 & 0 \\ 0 & 0 & 2/3 \end{pmatrix} M a^2$$

with axis of rotation at origin.

- Now we use parallel axis theorem to find \underline{I}_{cm} at CM:

$$I_{ij}^{cm} = I_{ij} - M(R_i R_j - R^2)$$



$$I_{xx}^{cm} = \frac{1}{6}Ma^2 - M((X_{cm}^2 + Y_{cm}^2) - X_{cm}^2)$$

$$= \frac{1}{6}Ma^2 - M(\frac{1}{9}a^2) = (\frac{3-2}{18})Ma^2 = \frac{1}{18}Ma^2$$

$$I_{yy}^{cm} = \frac{1}{2}Ma^2 - M((X_{cm}^2 + Y_{cm}^2) - Y_{cm}^2)$$

$$= \frac{1}{2}Ma^2 - \frac{4}{9}Ma^2 = \frac{9-8}{18}Ma^2 = \frac{1}{18}Ma^2$$

$$I_{xy}^{cm} = -\frac{1}{4}Ma^2 + M(X_{cm}Y_{cm})$$

$$= -\frac{1}{4}Ma^2 + \frac{2}{9}Ma^2 = (\frac{-9+8}{36})Ma^2 = -\frac{1}{36}Ma^2$$

$$I_{zz}^{cm} = \frac{2}{3}Ma^2 - M(X_{cm}^2 + Y_{cm}^2)$$

$$= \frac{2}{3}Ma^2 - M(\frac{1}{9}a^2 + \frac{4}{9}a^2) = (\frac{6-5}{9})Ma^2 = \frac{1}{9}Ma^2$$

$$\Rightarrow \underline{\underline{I}}^{cm} = \begin{pmatrix} \frac{1}{18} & -\frac{1}{36} & 0 \\ -\frac{1}{36} & \frac{1}{18} & 0 \\ 0 & 0 & \frac{1}{9} \end{pmatrix} Ma^2$$

To get to the principle moments of inertia and principle axes, we need to find the eigenvalues and eigenvectors of $\underline{\underline{I}}^{cm}$.

λ:

$$\det(\underline{\underline{I}}^{cm} - \lambda \underline{\underline{1}}) = 0$$

$$(\frac{1}{9} - \lambda) \left((\frac{1}{18} - \lambda)^2 - (\frac{1}{36})^2 \right) = 0$$

$$\lambda = \frac{1}{9} \quad \&$$

$$\frac{1}{18} - \lambda = \pm \frac{1}{36}$$

$$\lambda = \frac{1}{18} \pm \frac{1}{36}$$

$$= \frac{3}{36} \quad \text{or} \quad \frac{1}{36}$$

$$\lambda = \frac{1}{12} \quad \text{or} \quad \frac{1}{36}$$

For $\lambda = \frac{1}{9}$,

$$\begin{pmatrix} -\frac{1}{18} & -\frac{1}{36} & 0 \\ -\frac{1}{36} & -\frac{1}{18} & 0 \\ 0 & 0 & 0 \end{pmatrix} \begin{pmatrix} x \\ y \\ z \end{pmatrix} = 0$$

$$\Rightarrow 2x + y = 0 \quad \text{and} \quad x + 2y = 0$$

$$\Rightarrow x = y = 0$$

$$\Rightarrow \text{eigenvector is } \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$

→ So, we have $I_3 = \frac{1}{9} M a^2$ & the principle axis is \hat{z} .

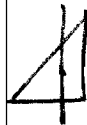
For $\lambda = \frac{1}{12}$, $I_1 = \frac{1}{12} M a^2$

$$\begin{pmatrix} \frac{1}{18} - \frac{1}{12} & -\frac{1}{36} & 0 \\ -\frac{1}{36} & \frac{1}{18} - \frac{1}{12} & 0 \\ 0 & 0 & \frac{1}{9} - \frac{1}{12} \end{pmatrix} \begin{pmatrix} x \\ y \\ z \end{pmatrix} = 0$$

$$\Rightarrow -\frac{1}{36}x - \frac{1}{36}y = 0 \quad \Rightarrow x + y = 0$$

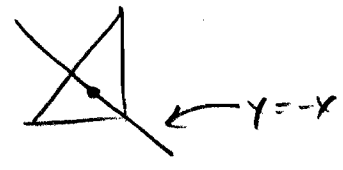
$$z = 0$$

$$\Rightarrow \begin{pmatrix} \frac{1}{\sqrt{2}} \\ -\frac{1}{\sqrt{2}} \\ 0 \end{pmatrix}$$



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→ So, we have $I_1 = \frac{1}{12} M a^2$ and the principle axis is



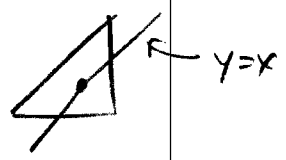
For $\lambda = \frac{1}{36}$, $I_2 = \frac{1}{36} M a^2$

$$\begin{pmatrix} \frac{1}{18} - \frac{1}{36} & -\frac{1}{36} & 0 \\ -\frac{1}{36} & \frac{1}{18} - \frac{1}{36} & 0 \\ 0 & 0 & \frac{1}{9} - \frac{1}{36} \end{pmatrix} \begin{pmatrix} x \\ y \\ z \end{pmatrix} = 0$$

$$\Rightarrow \frac{1}{36} x - \frac{1}{36} y = 0$$
$$z = 0$$

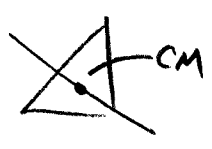
$$\Rightarrow x - y = 0 \Rightarrow \begin{pmatrix} 1/\sqrt{2} \\ 1/\sqrt{2} \\ 0 \end{pmatrix}$$

→ So, we have $I_2 = \frac{1}{36} M a^2$, and



Summary :

$$I_1 = \frac{1}{12} M a^2 \rightarrow$$



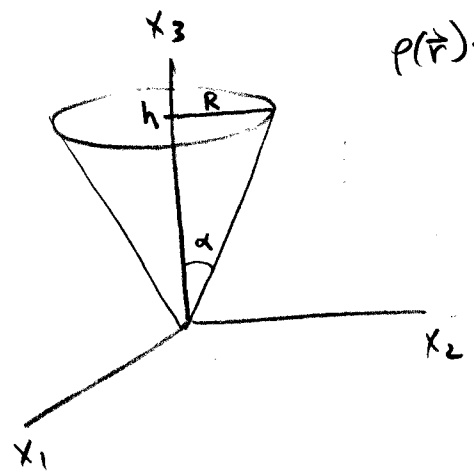
$$I_2 = \frac{1}{36} M a^2 \rightarrow$$



$$I_3 = \frac{1}{9} M a^2 \rightarrow$$



MT 11.2 :



$\rho(\vec{r}) = \rho$ constant mass density

$$I_{ij} = \int \rho(\vec{r}) [\delta_{ij} x_k x_k - x_i x_j] dv$$

- For convenience, we will use cylindrical coordinates.

$$I_{11} = \rho \int (x_2^2 + x_3^2) dx_1 dx_2 dx_3$$

← in rect. coordinates

$$= \rho \int_0^h \int_0^{2\pi} \int_0^{\frac{x_3 R}{h}} r^2 \cos^2 \theta + x_3^2 r dr d\theta dx_3$$

← in cylindrical coordinates

Note: $\int_0^{2\pi} \cos^2 \theta d\theta = \int_0^{2\pi} \left(\frac{1}{2} + \frac{\cos 2\theta}{2} \right) d\theta = \pi - \sin 2\theta \Big|_0^{2\pi} = \pi$

- so, doing the θ integral, we have,

$$\begin{aligned} I_{11} &= \rho \int_0^h \int_0^{\frac{x_3 R}{h}} [\pi r^2 + 2\pi x_3^2] r dr dx_3 \\ &= \rho \int_0^h \left[\pi \frac{r^4}{4} \Big|_0^{\frac{x_3 R}{h}} + 2\pi x_3^2 \frac{r^2}{2} \Big|_0^{\frac{x_3 R}{h}} \right] dx_3 \\ &= \rho \pi \int_0^h \left[\frac{1}{4} \left(\frac{R}{h} \right)^4 x_3^4 + \left(\frac{R}{h} \right)^2 x_3^4 \right] dx_3 \\ &= \rho \pi \left[\frac{1}{4} \left(\frac{R}{h} \right)^4 \frac{h^5}{5} + \left(\frac{R}{h} \right)^2 \frac{h^5}{5} \right] \\ &= \rho \pi \left(\frac{R}{h} \right)^2 \frac{h^5}{5} \left(\frac{1}{4} \left(\frac{R}{h} \right)^2 + 1 \right) \end{aligned}$$

(2)

- Now, we calculate the total mass M for the cone.

$$M = \int_0^h \int_0^{2\pi} \int_0^{x_3 R/h} \rho r dr d\theta dx_3$$

$$= 2\pi \rho \int_0^h \left(\frac{R}{h}\right)^2 \frac{x_3^2}{2} dx_3$$

$$\underline{M = \pi \rho \left(\frac{R}{h}\right)^2 \frac{h^3}{3}}$$

- So, $I_{11} = M \left(\frac{3}{5} h^2\right) \left(\frac{1}{4} \left(\frac{R}{h}\right)^2 + 1\right) = \frac{3}{20} M (R^2 + 4h^2)$

- The cone is symmetric with respect to the x_1 - x_2 plane,

So,

$$\underline{I_{22} = I_{11} = \frac{3}{20} M (R^2 + 4h^2)}$$

- Now, we calculate I_{33} :

$$I_{33} = \rho \int (x_1^2 + x_2^2) dx_1 dx_2 dx_3$$

$$= \rho \int_0^h \int_0^{2\pi} \int_0^{x_3 R/h} r^2 r dr d\theta dx_3$$

$$= 2\pi \rho \int_0^h \frac{1}{4} \left(\frac{R}{h}\right)^4 x_3^4 dx_3$$

$$= 2\pi \rho \left(\frac{1}{4} \left(\frac{R}{h}\right)^4\right) \frac{h^5}{5}$$

$$= \pi \rho \left(\frac{R}{h}\right)^2 \frac{h^3}{3} \left(\frac{1}{2} \left(\frac{R}{h}\right)^2 \cdot h^2 \left(\frac{3}{5}\right)\right)$$

$$I_{33} = \frac{3}{10} M R^2$$

3

$$I_{12} = \rho \int (-x_1 x_2) dx_1 dx_2 dx_3$$

$$= -\rho \int_0^h \int_0^{2\pi} \int_0^{\frac{x_3 R}{h}} r^2 \cos\theta \sin\theta r dr d\theta dx_3$$

$$= 0 \quad \text{since} \quad \int_0^{2\pi} \cos\theta \sin\theta d\theta = 0$$

$$I_{13} = \rho \int (-x_1 x_3) dx_1 dx_2 dx_3$$

$$= -\rho \int_0^h \int_0^{2\pi} \int_0^{\frac{x_3 R}{h}} r \cos\theta x_3 r dr d\theta dx_3$$

$$= 0 \quad \text{since} \quad \int_0^{2\pi} \cos\theta d\theta = 0$$

$$I_{23} = \rho \int (-x_2 x_3) dx_1 dx_2 dx_3$$

$$= -\rho \int_0^h \int_0^{2\pi} \int_0^{\frac{x_3 R}{h}} r \sin\theta x_3 r dr d\theta dx_3$$

$$= 0 \quad \text{since} \quad \int_0^{2\pi} \sin\theta d\theta = 0$$

So, we have

$$I_{zz} = \begin{pmatrix} \frac{3}{20} M (R^2 + 4h^2) & 0 & 0 \\ 0 & \frac{3}{20} M (R^2 + 4h^2) & 0 \\ 0 & 0 & \frac{3}{10} MR^2 \end{pmatrix}$$

with the cone upright along the x_2 axis and the apex at the origin.

- Now, we calculate the CM:

(4)

$$M X_1^{cm} = \int \rho x_1 dx_1 dx_2 dx_3$$
$$= \rho \int_0^h \int_0^{2\pi} \int_0^{x_3 R/h} r \cos \theta r dr d\theta dx_3 = 0$$

$$\underline{M X_2^{cm} = X_1^{cm} = 0}$$

$$M X_3^{cm} = \int \rho x_3 dx_1 dx_2 dx_3$$
$$= \rho \int_0^h \int_0^{2\pi} \int_0^{x_3 R/h} x_3 r dr d\theta dx_3$$
$$= 2\pi \rho \int_0^h x_3 \frac{1}{2} \left(\frac{R}{h}\right)^2 x_3^2 dx_3$$
$$= \pi \rho \left(\frac{R}{h}\right)^2 \frac{h^4}{4}$$

$$M X_3^{cm} = \frac{\pi \rho \left(\frac{R}{h}\right)^2 \frac{h^4}{4}}{M} \left(\frac{3}{4}h\right)$$

$$\underline{X_3^{cm} = \frac{3}{4}h}$$

- Now, we use the parallel axis theorem to express the I_{cm} with the origin at the CM.

$$I_{ij}^{CM} = I_{ij} - M \delta_{ij} a^2 - a_i a_j$$

$$\vec{a} = (0, 0, \frac{3}{4}h)$$

$$I_{11}^{CM} = I_{11} - M \frac{9}{16} h^2$$

$$= \frac{3}{20} M (R^2 + 4h^2) - \frac{9}{16} M h^2$$

$$= \frac{3}{20} M R^2 + \frac{3}{5} M h^2 - \frac{9}{16} M h^2$$

$$= \frac{3}{20} M R^2 + M h^2 \left(\frac{48}{80} - \frac{45}{80} \right)$$

$$I_{11}^{CM} = \frac{3}{20} M \left(R^2 + \frac{h^2}{4} \right)$$

$$I_{22}^{CM} = I_{11}^{CM}$$

$$I_{33}^{CM} = I_{33} - M \left(\frac{9}{16} h^2 - \frac{9}{16} h^2 \right) = I_{33} = \frac{3}{10} M R^2$$

So,

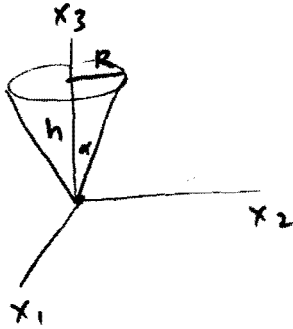
$$I_{\hat{x}}^{CM} = \begin{pmatrix} \frac{3}{20} M \left(R^2 + \frac{h^2}{4} \right) & 0 & 0 \\ 0 & \frac{3}{20} M \left(R^2 + \frac{h^2}{4} \right) & 0 \\ 0 & 0 & \frac{3}{10} M R^2 \end{pmatrix}$$

Since this is diagonalized already, the diagonal elements are the principle moments.

5.17

①

From MT 11.2, we have computed the moments of inertia of a cone rotating about its symmetry axis and with the origin of a coordinate system at the apex of the cone as shown below:



$$\tan \alpha = \frac{R}{h}, \quad \rho - \text{uniform density}$$

$$M(\text{total mass}) = \pi \rho \left(\frac{R}{h}\right)^2 \frac{h^3}{3}$$

$$\underline{I} = \begin{pmatrix} \frac{3}{20} M(R^2 + 4h^2) & 0 & 0 \\ 0 & \frac{3}{20} M(R^2 + 4h^2) & 0 \\ 0 & 0 & \frac{3}{10} MR^2 \end{pmatrix}$$

For this problem we will have this cone rolling on its side without slipping on a uniform horizontal plane. This is in the space axes. Let call this (x'_1, x'_2, x'_3) . The (x_1, x_2, x_3) coordinate system in which we calculated \underline{I} is the body axes. These two coordinate systems can be related by the Euler's angles.

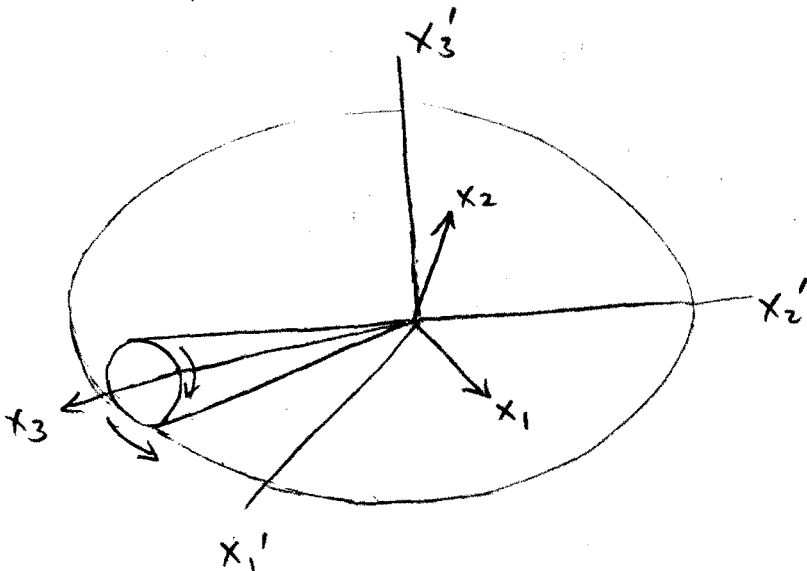
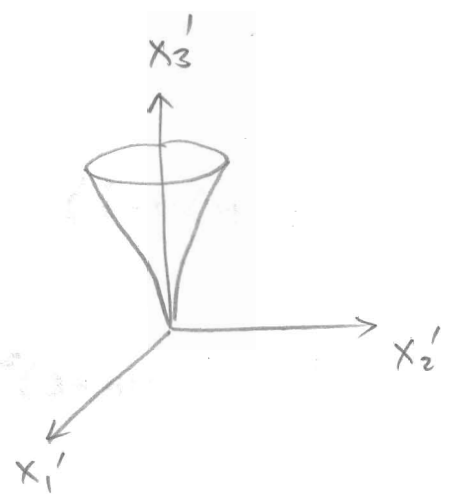


Fig. 1

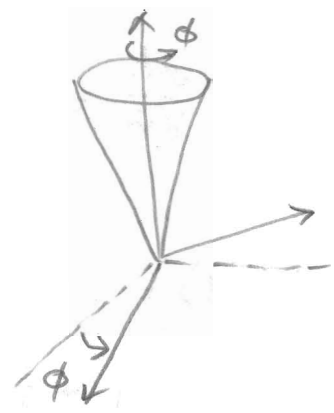
Cone rolling on its side on a horizontal plane.

(x'_1, x'_2, x'_3) - space axes
 (x_1, x_2, x_3) - body axes



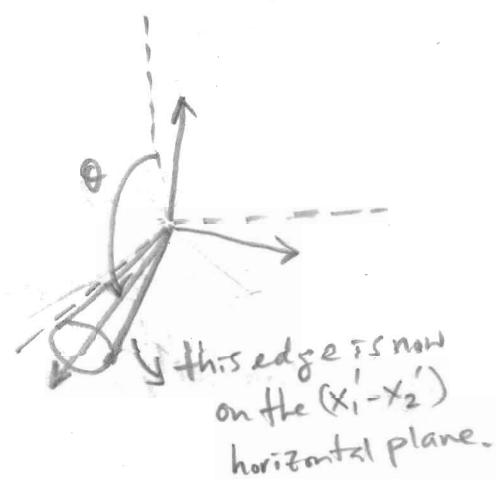
Euler Angle ϕ

→



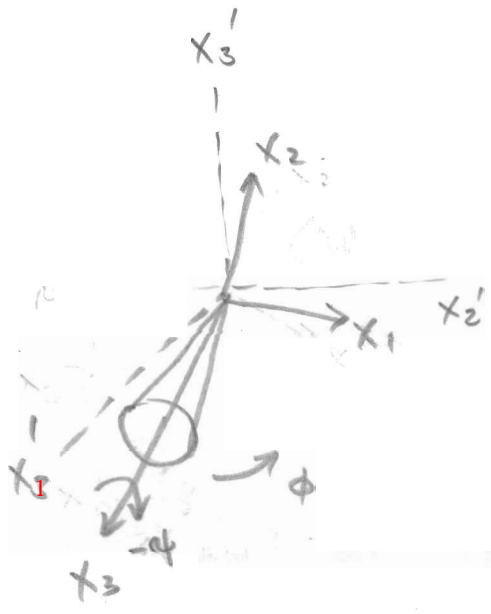
Euler Angle θ

↓



Euler Angle ψ

←



- We are given that the cone rolls around (without slipping) the x_3' -axis (space axis) in a time τ .

→ We will define $\omega_0 \equiv$ frequency of rolling around the x_3' -axis

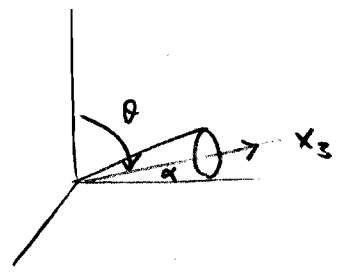
where
$$\omega_0 = \frac{2\pi}{\tau} = \dot{\phi}$$

→ And, define $\omega_s \equiv$ frequency of spinning of the cone around its x_3 body axis.

With the orientation shown in Fig. 1,

$\omega_s = -\dot{\psi}$

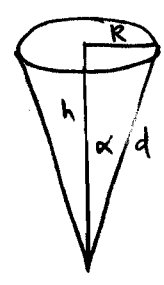
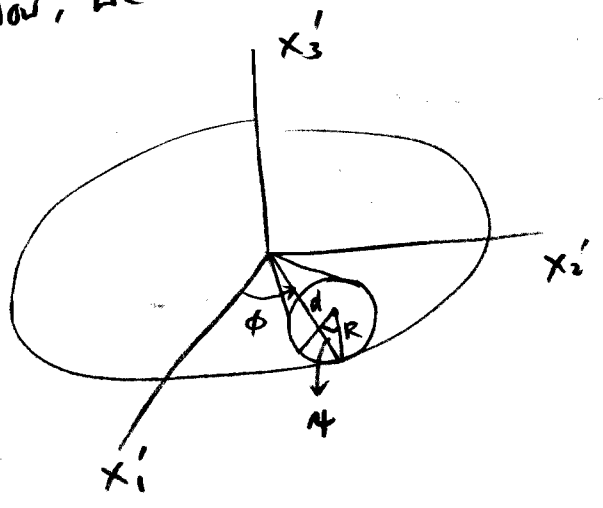
- Together with the geometrical constraint of the cone laying on its edge,



We have the following equations for the Euler's angles:

(*) $\left\{ \begin{array}{l} \theta = \frac{\pi}{2} - \alpha \\ \phi = \omega_0 t \\ \psi = -\omega_s t \end{array} \right.$

- Now, we consider the no slip condition:



$R/d = \sin \alpha$

for rolling without slipping, we need to have the following distances to be equal:

$$|\phi d| = |\omega R|$$

For constant rolling speeds, we have

$$(\omega_0 t) d = (\omega_s t) R$$

$$\omega_s = \frac{\omega_0}{\sin \alpha}$$

So, the magnitude of the two angular speeds are related.

- With $\hat{\phi}$ well defined, we can write down the components of the angular velocity vector $\vec{\omega}$ for the core's motion in the body axes using the equations on p. 174 (Eq. 4.87):

$$\begin{aligned} \omega_1 &= \dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos \psi \\ \omega_2 &= \dot{\phi} \sin \theta \cos \psi - \dot{\theta} \sin \psi \\ \omega_3 &= \dot{\phi} \cos \theta + \dot{\psi} \end{aligned}$$

$$\vec{\omega} = \omega_1 \hat{x}_1 + \omega_2 \hat{x}_2 + \omega_3 \hat{x}_3$$

- Substituting what we have for ϕ, θ, ψ & their derivatives, we have

$$\begin{aligned} \omega_1 &= \omega_0 \sin\left(\frac{\pi}{2} - \alpha\right) \sin(-\omega_s t) + 0 \\ \omega_2 &= \omega_0 \sin\left(\frac{\pi}{2} - \alpha\right) \cos(-\omega_s t) + 0 \\ \omega_3 &= \omega_0 \cos\left(\frac{\pi}{2} - \alpha\right) - \omega_s \end{aligned}$$

$$\Rightarrow \begin{cases} \omega_1 = -\omega_0 \cos \alpha \sin(\omega_0 t) \\ \omega_2 = \omega_0 \cos \alpha \cos(\omega_0 t) \\ \omega_3 = \omega_0 \sin \alpha - \frac{\omega_0}{\sin \alpha} \end{cases}$$

$$= \omega_0 \left(\frac{\sin^2 \alpha - 1}{\sin \alpha} \right)$$

$$= -\omega_0 \frac{\cos^2 \alpha}{\sin \alpha} = -\omega_0 \frac{\cos \alpha}{\tan \alpha}$$

- In the body axes, the moments of inertia are given by

$$I_{11} = I_{22} = \frac{3}{20} M (R^2 + 4h^2)$$

$$\& I_{12} = I_{13} = I_{23} = 0$$

$$I_{33} = \frac{3}{10} MR^2$$

$$\begin{aligned} \text{- Thus, } T &= \frac{1}{2} I_{11} \omega_1^2 + \frac{1}{2} I_{22} \omega_2^2 + \frac{1}{2} I_{33} \omega_3^2 \\ &= \frac{1}{2} \left(\frac{3}{20} M (R^2 + 4h^2) \right) (\omega_1^2 + \omega_2^2) + \frac{1}{2} \left(\frac{3}{10} MR^2 \right) \omega_3^2 \\ &= \frac{3}{40} M (R^2 + 4h^2) \left(\omega_0^2 \cos^2 \alpha (\sin^2(\omega_0 t) + \cos^2(\omega_0 t)) \right) \\ &\quad + \frac{3}{20} MR^2 \omega_0^2 \frac{\cos^2 \alpha}{\tan^2 \alpha} \end{aligned}$$

$$\tan \alpha = \frac{R}{h}$$

$$= \frac{3}{40} M \omega_0^2 \cos^2 \alpha \left(R^2 + 4h^2 + 2R^2 \frac{h^2}{R^2} \right)$$

$$T = \frac{3}{40} M \omega_0^2 \cos^2 \alpha (R^2 + 6h^2)$$

Now, we calculate the angular momentum \vec{L} with respect to the body axes :

$$L_1 = I_{11} \omega_1 = \frac{3}{20} M(R^2 + 4h^2) (-\omega_0 \cos \alpha \sin(\omega_0 t))$$

$$= -\frac{3}{20} M \omega_0 \cos \alpha (R^2 + 4h^2) \sin\left(\frac{\omega_0}{\sin \alpha} t\right)$$

$$L_2 = I_{22} \omega_2 = \frac{3}{20} M(R^2 + 4h^2) (\omega_0 \cos \alpha \cos(\omega_0 t))$$

$$= \frac{3}{20} M \omega_0 \cos \alpha (R^2 + 4h^2) \cos\left(\frac{\omega_0}{\sin \alpha} t\right)$$

$$L_3 = I_{33} \omega_3 = \frac{3}{10} M R^2 \left(-\omega_0 \frac{\cos \alpha}{\tan \alpha}\right)$$

$$= -\frac{3}{10} M \omega_0 \cos \alpha R^2 \left(\frac{h}{R}\right)$$

$$= -\frac{3}{10} M \omega_0 \cos \alpha R h$$

\vec{L} can be expressed in the space axes using the inverse Euler transform (Eq. 4-47) on p. 153.

5.26

For a symmetric top without external torques, we have ①

$$I_1 = I_2 \text{ and } \vec{N} = 0.$$

So, we have the following Euler's Equation:

$$\left\{ \begin{array}{l} I_1 \dot{\omega}_1 = (I_1 - I_3) \omega_2 \omega_3 \quad (1) \\ I_1 \dot{\omega}_2 = (I_3 - I_1) \omega_3 \omega_1 \quad (2) \\ I_3 \dot{\omega}_3 = 0 \quad (3) \end{array} \right.$$

③ $\Rightarrow \omega_3 = \text{const}$

For convenience, let $\Omega = \frac{I_3 - I_1}{I_1} \omega_3$ or $\omega_3 = \frac{I_1}{I_3 - I_1} \Omega$

Then, we can write ① & ② as

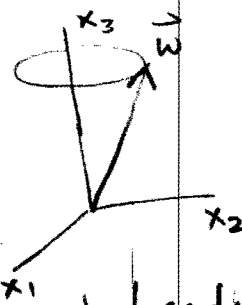
$$\left\{ \begin{array}{l} \dot{\omega}_1 = -\Omega \omega_2 \\ \dot{\omega}_2 = \Omega \omega_1 \end{array} \right.$$

The solution for this simple linear ODE is:

$$\omega_1 = A \cos(\Omega t + \delta)$$

$$\omega_2 = A \sin(\Omega t + \delta)$$

where A & δ will be determined by initial conditions.



$\vec{\omega}$ precesses with constant angular velocity Ω about x_3 .

Equation 4.87 relates the components of the angular velocity vector $\vec{\omega}$ in the body axes to the Euler's angles,

$\phi, \theta, \psi =$

$$\left\{ \begin{aligned} \omega_1 &= \dot{\phi} \sin\theta \sin\psi + \dot{\theta} \cos\psi = A \cos(\omega t + \delta) & (1)' \\ \omega_2 &= \dot{\phi} \sin\theta \cos\psi - \dot{\theta} \sin\psi = A \sin(\omega t + \delta) & (2)' \\ \omega_3 &= \dot{\phi} \cos\theta + \dot{\psi} = \left(\frac{I_1}{I_3 - I_1}\right) \Omega & (3)' \end{aligned} \right.$$

(1)' $\sin\psi +$ (2)' $\cos\psi =$

$$\begin{aligned} &\dot{\phi} \sin\theta \sin^2\psi + \dot{\theta} \cancel{\sin\psi} \cos\psi = A \cos(\omega t + \delta) \sin\psi \\ &+ \dot{\phi} \sin\theta \cos^2\psi - \dot{\theta} \cancel{\sin\psi} \cos\psi = + A \sin(\omega t + \delta) \cos\psi \end{aligned}$$

$$\Rightarrow \dot{\phi} \sin\theta = A \sin(\omega t + \delta + \psi) \quad (4)'$$

(1)' $\cos\psi -$ (2)' $\sin\psi =$

$$\begin{aligned} &\dot{\phi} \cancel{\sin\theta} \sin\psi \cos\psi + \dot{\theta} \cos^2\psi = A \cos(\omega t + \delta) \cos\psi \\ &- \dot{\phi} \cancel{\sin\theta} \sin\psi \cos\psi + \dot{\theta} \sin^2\psi = - A \sin(\omega t + \delta) \sin\psi \end{aligned}$$

$$\Rightarrow \dot{\theta} = A \cos(\omega t + \delta + \psi) \quad (5)'$$

For uniform precession, we need to have

(3)

$$\dot{\theta} = 0 \quad (\text{no nutation})$$

$$\dot{\phi} = \text{const} \quad (\text{constant speed})$$

$$(5)' \Rightarrow \dot{\theta} = A \cos(\Omega t + \delta + \psi) = 0$$

$$\Rightarrow \Omega t + \delta + \psi = (n + \frac{1}{2})\pi, \quad n = 0, 1, 2, \dots$$

Taking $n=0$ and calling $\psi_0 = -\delta + \pi/2$, we have

$$\psi = -\Omega t + \psi_0 \quad (*1)$$

$$\text{and } \dot{\theta} = 0 \Rightarrow \underline{\theta = \theta_0} \quad (*2)$$

— Using (*1) & (*2), (4)' gives

$$\dot{\phi} = A \sin(\pi/2) / \sin\theta_0 = \text{const.} \quad (*3) \quad (\text{as is required by uniform precession}).$$

— To solve for the constant A , we can substitute

(*1), (*2), & (*3) back into (3)', we have,

$$\omega_3 = A \frac{\cos\theta_0}{\sin\theta_0} - \Omega$$

$$(\omega_3 + \Omega) \tan \theta_0 = A$$

$$\left(\frac{I_1 + I_3 - I_1}{I_3 - I_1} \right) \Omega \tan \theta_0 = A$$

$$A = \Omega \tan \theta_0 \left(\frac{I_3}{I_3 - I_1} \right)$$

So, $\dot{\phi} = \Omega \frac{\tan \theta_0}{\sin \theta_0} \left(\frac{I_3}{I_3 - I_1} \right)$

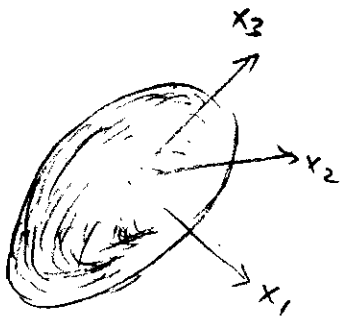
$\Rightarrow \phi = \frac{\Omega}{\cos \theta_0} \left(\frac{I_3}{I_3 - I_1} \right) t + \phi_0$ (*)

Summary :

$$\left\{ \begin{array}{l} \phi = \frac{\Omega}{\cos \theta_0} \left(\frac{I_3}{I_3 - I_1} \right) t + \phi_0 \\ \theta = \theta_0 \\ \psi = -\Omega t + \psi_0 \end{array} \right.$$

Problem 3 :

①



$$I_1 > I_2 > I_3$$

let $\vec{\omega} = (\omega_1, \omega_2, \omega_3)$ be the angular velocity of the rotating body. We need to consider three cases:

Case 1: rotating almost directly along x_1 .

If stable, small deviations away from x_1 will not grow in time.

To check for this, we can consider the time evolution of $\vec{\omega}$ using the Euler's Equations:

$$\begin{cases} I_1 \dot{\omega}_1 - \omega_2 \omega_3 (I_2 - I_3) = 0 & \textcircled{1} \\ I_2 \dot{\omega}_2 - \omega_3 \omega_1 (I_2 - I_1) = 0 & \textcircled{2} \\ I_3 \dot{\omega}_3 - \omega_1 \omega_2 (I_1 - I_2) = 0 & \textcircled{3} \end{cases}$$

- In this case, we have set initially $\omega_1 \gg \omega_2, \omega_3$.

- Since there is no torque acting on the system,

\vec{L} is fixed and $|\vec{\omega}|$ is also fixed.

(2)

- $|\vec{\omega}| = (\omega_1^2 + \omega_2^2 + \omega_3^2)^{1/2} = \text{const} \quad (*)$

- Since $\omega_1 \gg \omega_2, \omega_3$, $(*)$ implies that

$\omega_1 \approx \text{const}$ to 1st approximation.

Putting this into our Euler's Equations = {

- $(*)$ does not give us any useful information
- it is just a consistency statement saying that $\dot{\omega}_1 \approx \dot{\omega}_2 \approx \dot{\omega}_3 \approx 0$.
- $\dot{\omega}_2$ & $\dot{\omega}_3$ are not necessarily small and we have to 1st order approximation, the following coupled equations:

$$\begin{cases} \dot{\omega}_2 = \frac{\omega_1(I_3 - I_1)}{I_2} \omega_3 \\ \dot{\omega}_3 = \frac{\omega_1(I_1 - I_2)}{I_3} \omega_2 \end{cases}$$

- Taking the derivative on both sides, we have

$$\begin{cases} \ddot{\omega}_2 = \frac{\omega_1(I_3 - I_1)}{I_2} \dot{\omega}_3 = \frac{\omega_1^2 (I_3 - I_1)(I_1 - I_2)}{I_2 I_3} \omega_2 \\ \ddot{\omega}_3 = \frac{\omega_1(I_1 - I_2)}{I_3} \dot{\omega}_2 = \frac{\omega_1^2 (I_1 - I_2)(I_3 - I_1)}{I_3 I_2} \omega_3 \end{cases}$$

- (3)
- Now, the two equations have decoupled and the constant coefficients on the RHS are the same:

$$\frac{\omega_1^2 (I_1 - I_2)(I_3 - I_1)}{I_2 I_3} < 0$$

which is negative since $I_1 > I_2 > I_3$!

- Thus, the time evolution for w_2, w_3 are described by a harmonic equation:

$$\ddot{w}_{2,3} = -\Omega^2 w_{2,3}, \quad \Omega^2 = \frac{\omega_1^2 (I_1 - I_2)(I_1 - I_3)}{I_2 I_3} > 0$$

with $w_{2,3}(t) \sim e^{\pm i\Omega t}$

- Since $w_{2,3}$ will not grow in time, spinning motion along x_1 will remain stable!

Case 2: $w_2 \gg w_1, w_3$

Following a similar argument as case 1, we can arrive at the following 2nd order ODEs

for w_1 & w_3 :

$$\left\{ \begin{aligned} \ddot{w}_1 &= \frac{w_2 (I_2 - I_3)}{I_1} & \ddot{w}_3 &= \frac{w_2^2 (I_2 - I_3)(I_1 - I_2)}{I_1 I_3} w_1 \\ \ddot{w}_3 &= \frac{w_2 (I_1 - I_2)}{I_3} & \ddot{w}_1 &= \frac{w_2^2 (I_1 - I_3)(I_2 - I_3)}{I_3 I_1} w_3 \end{aligned} \right.$$

- Since $I_1 > I_2 > I_3$,

$$Y^2 = \frac{w_2^2 (I_2 - I_3)(I_1 - I_2)}{I_1 I_3} > 0$$

$$\Rightarrow w_{1,3}(t) \sim e^{\pm \lambda t}$$

⇒ Therefore, $w_{1,3}(t)$ might grow in time and spinning initially along x_2 is not stable!

Lastly case 3 : $w_3 \gg w_1, w_2$

In this case, we will have

$$\ddot{w}_1 = \frac{w_3^2 (I_2 - I_3)(I_3 - I_1)}{I_1 I_2} w_1$$

$$\ddot{w}_2 = \frac{w_3^2 (I_3 - I_1)(I_2 - I_3)}{I_1 I_2} w_2$$

$$\text{Since, } \frac{\omega_3^2 (I_2^+ - I_3)(I_3 - I_1^-)}{I_1 I_2} = -\omega'^2 < 0$$

(5)

$\omega_{1,2}(t) \sim e^{\pm i\omega' t}$ are harmonic
and spinning along x_3 initially will be
stable!