

)

$$(\mathbf{A} \times \mathbf{B})_j = \sum_{lm} \epsilon_{jlm} A_l B_m$$

$$(\mathbf{C} \times \mathbf{D})_k = \sum_{rs} \epsilon_{krs} C_r D_s$$

$$\begin{aligned}[(\mathbf{A} \times \mathbf{B}) \times (\mathbf{C} \times \mathbf{D})]_i &= \sum_{jk} \epsilon_{ijk} \left(\sum_{lm} \epsilon_{jlm} A_l B_m \right) \left(\sum_{rs} \epsilon_{krs} C_r D_s \right) \\&= \sum_{jklmrs} \epsilon_{ijk} \epsilon_{jlm} \epsilon_{krs} A_l B_m C_r D_s \\&= \sum_{jlmrs} \epsilon_{jlm} \left(\sum_k \epsilon_{ijk} \epsilon_{rsk} \right) A_l B_m C_r D_s \\&= \sum_{jlmrs} \epsilon_{jlm} (\delta_{ir} \delta_{js} - \delta_{is} \delta_{jr}) A_l B_m C_r D_s \\&= \sum_{jlm} \epsilon_{jlm} (A_l B_m C_r D_j - A_l B_m D_r C_j) \\&= \left(\sum_{jlm} \epsilon_{jlm} D_j A_l B_m \right) C_i - \left(\sum_{jlm} \epsilon_{jlm} C_j A_l B_m \right) D_i \\&= (\mathbf{ABD})\mathbf{C}_i - (\mathbf{ABC})\mathbf{D}_i\end{aligned}$$

Therefore,

$$[(\mathbf{A} \times \mathbf{B}) \times (\mathbf{C} \times \mathbf{D})] = (\mathbf{ABD})\mathbf{C} - (\mathbf{ABC})\mathbf{D}$$

2)

First $(\mathbf{B} \times \mathbf{C})_i = \sum_{jk} \varepsilon_{ijk} B_j C_k$. Then,

$$\begin{aligned}
 [\mathbf{A} \times (\mathbf{B} \times \mathbf{C})]_t &= \sum_{mn} \varepsilon_{tmm} A_m (\mathbf{B} \times \mathbf{C})_n = \sum_{mn} \varepsilon_{tmm} A_m \sum_{jk} \varepsilon_{njk} B_j C_k \\
 &= \sum_{jkm} \varepsilon_{tmm} \varepsilon_{njk} A_m B_j C_k = \sum_{jkm} \varepsilon_{tmm} \varepsilon_{jkn} A_m B_j C_k \\
 &= \sum_{jkm} \left(\sum_n \varepsilon_{lmn} \varepsilon_{jkn} \right) A_m B_j C_k \\
 &= \sum_{jkm} (\delta_{jl} \delta_{km} - \delta_{kl} \delta_{jm}) A_m B_j C_k \\
 &= \sum_m A_m B_t C_m - \sum_m A_m B_m C_t = B_t \left(\sum_m A_m C_m \right) - C_t \left(\sum_m A_m B_m \right) \\
 &= (\mathbf{A} \cdot \mathbf{C}) B_t - (\mathbf{A} \cdot \mathbf{B}) C_t
 \end{aligned}$$

Therefore,

$$\boxed{\mathbf{A} \times (\mathbf{B} \times \mathbf{C}) = (\mathbf{A} \cdot \mathbf{C}) \mathbf{B} - (\mathbf{A} \cdot \mathbf{B}) \mathbf{C}}$$

(9)

4.15

$\vec{\omega} = \vec{\omega}_\phi + \vec{\omega}_\theta + \vec{\omega}_\psi$. Write its components out in the space set of axes in terms of the Euler angles.

$\vec{\omega}_\phi$, $\vec{\omega}_\theta$, and $\vec{\omega}_\psi$ are the three successive infinitesimal rotations with angular velocities

$$\vec{\omega}_\phi = \begin{pmatrix} 0 \\ 0 \\ \dot{\phi} \end{pmatrix}, \quad \vec{\omega}_\theta = \begin{pmatrix} 0 \\ \dot{\theta} \\ 0 \end{pmatrix}, \quad \vec{\omega}_\psi = \begin{pmatrix} 0 \\ 0 \\ \dot{\psi} \end{pmatrix}$$

along \hat{z} along the line
of nodes \hat{z} along \hat{z}'

① $\vec{\omega}_\phi$ is already along the space axis, no change is needed.

② $\vec{\omega}_\theta$ is along the line of nodes. To express it back to the space axes, rotate it by $-\phi$ using the D-matrix (inverse of $D \cdot D^{-1}$):

$$\begin{pmatrix} \cos\phi & -\sin\phi & 0 \\ \sin\phi & \cos\phi & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 0 \\ \dot{\theta} \\ 0 \end{pmatrix} = \begin{pmatrix} 0 \cos\phi \\ 0 \sin\phi \\ 0 \end{pmatrix} = \vec{\omega}_\theta$$

Note: To get from the \hat{z} set of intermediate axes back to \hat{x} , we formally need $(CD)^{-1} = D^T C^{-1}$. However, since C^{-1} simply rotates x-component of a vector, it does not affect $\vec{\omega}_\theta$. It suffices to use D^T only.

(2)

- ③ \vec{w}_4 is along the body axes E' . To express it back to the space axes, rotate it by the inverse $(BCD)^{-1} = D^T C^{-1} B^{-1}$. Again, since E' rotates the z-component of a vector, it suffices to consider $D^T C^T$ only:

$$\begin{aligned} & \begin{pmatrix} \cos\phi & -\sin\phi & 0 \\ \sin\phi & \cos\phi & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta & -\sin\theta \\ 0 & \sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} 0 \\ 0 \\ \dot{\psi} \end{pmatrix} \\ &= \begin{pmatrix} \cos\phi & -\sin\phi \cos\theta & \sin\phi \sin\theta \\ \sin\phi & \cos\phi \cos\theta & -\sin\phi \sin\theta \\ 0 & \sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} 0 \\ 0 \\ \dot{\psi} \end{pmatrix} \\ &= \begin{pmatrix} i\dot{\psi} \sin\phi \sin\theta \\ -i\dot{\psi} \cos\phi \sin\theta \\ i\dot{\psi} \cos\theta \end{pmatrix} = (\vec{w}_4)_S \end{aligned}$$

Thus, in the space axes,

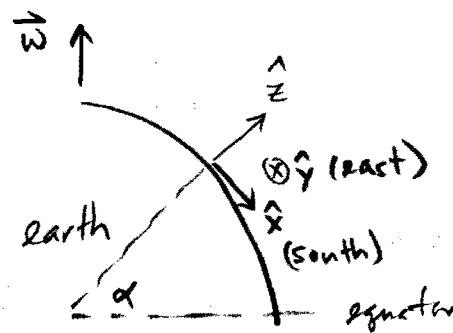
$$(\vec{w})_S = \begin{pmatrix} \dot{\theta} \cos\phi + i\dot{\psi} \sin\phi \sin\theta \\ \dot{\theta} \sin\phi - i\dot{\psi} \cos\phi \sin\theta \\ i\dot{\psi} \cos\theta \end{pmatrix}$$

10

4.21

11

Here is the geometry of the coordinate system:



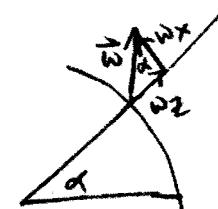
- Considering the Coriolis effect alone (ignoring the centrifugal force), the equation of motion for a particle moving with velocity \vec{v} is given by:

$$-2m(\vec{\omega} \times \vec{v}) - mg\hat{z} = m\vec{a}$$

$$\Rightarrow \vec{v} = -2(\vec{\omega} \times \vec{v}) - g\hat{z} \quad (1)$$

- For the coordinates chosen, we have

$$\vec{\omega} = -\omega \cos \alpha \hat{x} + \omega \sin \alpha \hat{z}$$



$$\text{- Then, } \vec{\omega} \times \vec{v} = \begin{pmatrix} \hat{x} & \hat{y} & \hat{z} \\ w_x & 0 & w_z \\ v_x & v_y & v_z \end{pmatrix}$$

$$= -w_z v_y \hat{x} + (w_z v_x - w_x v_z) \hat{y} + w_x v_y \hat{z}$$

So that, in component form, we can write :

$$\left\{ \begin{array}{l} \ddot{v}_x = 2\omega \sin \alpha v_y \\ \ddot{v}_y = -2\omega (\cos \alpha v_z + \sin \alpha v_x) \\ \ddot{v}_z = 2\omega \cos \alpha v_y - g \end{array} \right. \quad (*)$$

- Obviously, to zero order in ω , we have the simple free fall situation with

$$\dot{v}_x^{(0)} = 0 ; \dot{v}_y^{(0)} = 0 ; \dot{v}_z^{(0)} = -g$$

Case 1: with the initial condition (throwing straight up with velocity $v(0) = v_0 \hat{z}$), we have

$$v_z^{(0)} = v_0 - gt$$

$$\text{and } z^{(0)}(t) = z_0 + v_0 t - \frac{1}{2}gt^2$$

$$\text{And, } v_x^{(0)} = 0 \quad \text{and} \quad v_y^{(0)} = 0$$

- Now, we will consider the perturbative solution upto $O(\omega)$ by substituting $\vec{v}^{(0)}$ into the RHS of $(*)$:

$$\dot{v}_x^{(1)} = 0$$

$$\dot{v}_y^{(1)} = -2\omega \cos \alpha (v_0 - gt)$$

$$\dot{v}_z^{(1)} = -g$$

(3)

- Up to $O(\omega)$, the Coriolis force affect only the y-dir only.

(From now on, I will suppress the order notation.)

$$v_y = -2\omega \cos \alpha (v_0 - gt) \quad \text{I.C.}$$

$$\text{so, } v_y = -2\omega \cos \alpha v_0 t + \frac{g}{2} \omega \cos \alpha t^2 \quad (v_y(0) = 0)$$

$$y = -\omega \cos \alpha v_0 t^2 + \frac{1}{3} \frac{g}{2} \omega \cos \alpha t^3 \quad (y(0) = 0)$$

$$v_z = v_0 - gt ; \quad z = v_0 t - \frac{1}{2} g t^2 \quad (z(0) = 0)$$

- The particle will reach its maximum height h when $v_z = 0$ at $t = \frac{v_0}{g}$ (top)

$$\therefore h = \frac{v_0^2}{g} - \frac{1}{2} g \left(\frac{v_0}{g} \right)^2 = \frac{v_0^2}{2g}$$

- It will fall back down to the ground at $0 = v_0 - \frac{1}{2} g t \Rightarrow t = \frac{2v_0}{g}$ (ground again)

- At this time, the y deflection is :

$$y = -\omega \cos \alpha v_0 \left(\frac{2v_0}{g} \right)^2 + \frac{1}{3} \frac{g}{2} \omega \cos \alpha \left(\frac{2v_0}{g} \right)^3$$

$$= \omega \cos \alpha \frac{v_0^3}{g^2} \left(-4 + \frac{8}{3} \right) = -\underbrace{\frac{4}{3} \omega \cos \alpha \frac{v_0^3}{g^2}}$$

The negative sign indicates that the deflection is to the west. (4)

Case 2 :

Now, we consider the other situation when the particle is released at h at rest instead:

$$\text{I.C. : } z(0) = h \quad , \quad v_z(0) = v_x(0) = v_y(0) = 0$$

From the z equation, we can calculate the time when it hits the ground.

$$z = z_0 + v_{z0}t - \frac{1}{2}gt^2$$

$$0 = h - \frac{1}{2}gt^2$$

$$t = \sqrt{\frac{2h}{g}}$$

The y -dir equation is,

$$\dot{v}_y = -w \cos \alpha v_z = -w \cos \alpha (-gt) \quad \left\{ \begin{array}{l} v_z = -8 \\ v_z = -8t \\ + 25^\circ \end{array} \right.$$

$$v_y = w \cos \alpha g t^2$$

$$Y = \frac{1}{3} w \cos \alpha g t^3$$

$$\text{At } t = \sqrt{\frac{2h}{g}}, \quad Y = \frac{1}{3} w \cos \alpha g \left(\frac{2h}{g}\right)^{3/2}$$

In terms of v_0 from the other result, we have

(5)

$$h = \frac{v_0^2}{2g}$$

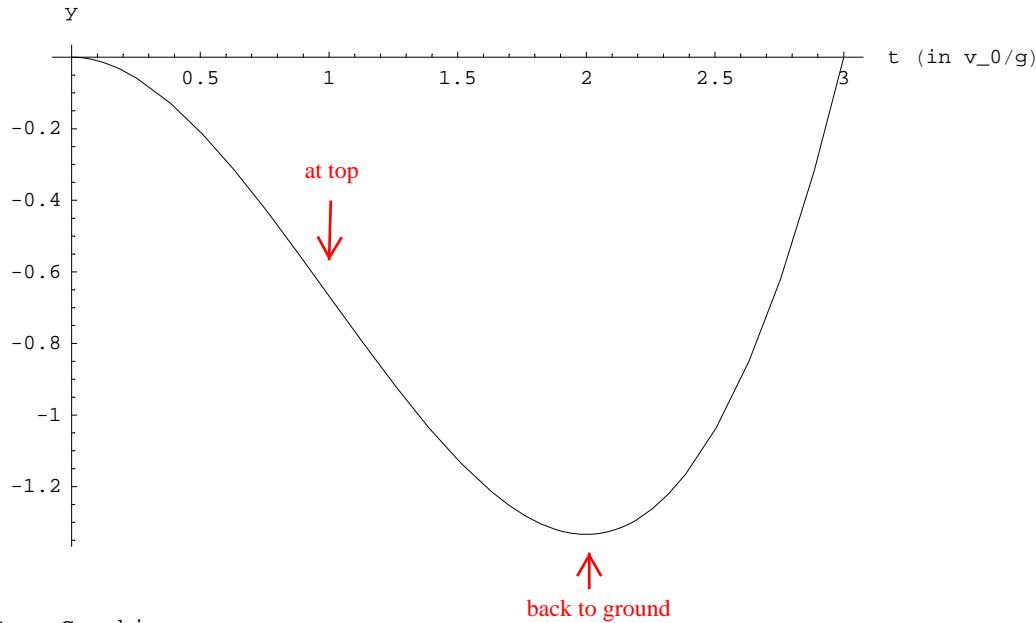
$$\text{so, } \frac{2h}{g} = \frac{v_0^2}{g^2}$$

$$\text{Then, } y = \frac{1}{3} N \cos \alpha g \left(\frac{v_0^2}{g^2} \right)^{3/2} = \frac{1}{3} N \cos \alpha g \frac{v_0^3}{g^{5/2}}$$
$$= \frac{1}{3} N \cos \alpha \frac{v_0^3}{g^{5/2}}$$

- The + sign indicates that the deflection is to the east!
And it has a magnitude which is $\frac{1}{4}$ of the other result.

$$\Rightarrow Y_{\text{case 1}} = -4 Y_{\text{case 2}}$$

```
In[19]:= Plot[{-t^2 + t^3/3}, {t, 0, 3}, AxesLabel -> {"t (in v_0/g)", "y"}]
```



```
Out[19]= - Graphics -
```

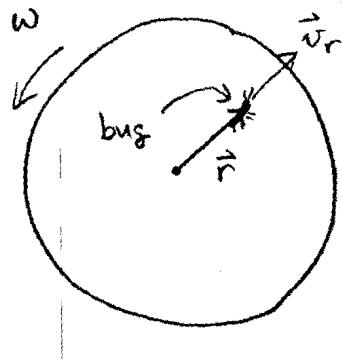
Goldstein P 4.21: Plot of the 1st order deflection in y with the ball initially thrown straight up vertically. Time is in units of v_0/g . The ball reaches the top of its trajectory when $t=v_0/g$ and it falls back down to the ground at $t=2v_0/g$. The deflection is initially westward ($-y$) and it becomes eastward ($+y$) only after $t>=3v_0/g$.

Note: In Case 1, v_y is non-zero (in fact still negative --> going westward) at the top of trajectory so that it continues to move westward as it comes back down while its v_y continues to decrease eventually going back to zero at ground level. However, as a result, the trajectory still has a westward deflection. In fact, it has its maximum westward deflection when it comes back to the ground (see above figure).

In Case 2, the particle is dropped from REST with ALL its velocity components equal to zero.

4.24

①



Given :

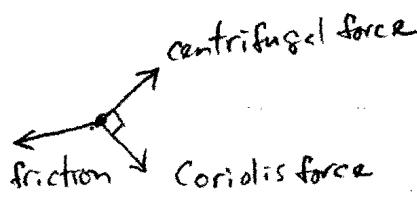
$\omega = 3.0 \text{ radian/s}$ (out of page)

$$\vec{v}_r = v_r \hat{r}, \quad v_r = 0.5 \text{ cm/s}$$

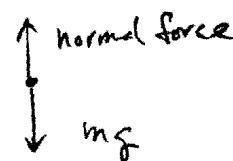
$$\mu = 0.30$$

- Free body diagram in the rotating frame:

Top View



Side View

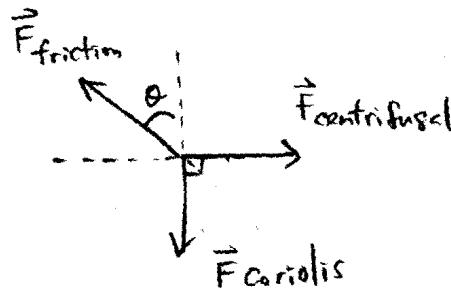


- In the rotating frame, we have the following equation of motion.

$$\vec{F}_{\text{eff}} = \vec{F} - 2m(\vec{\omega} \times \vec{v}_r) - m\vec{\omega} \times (\vec{\omega} \times \vec{r}) = m\vec{a}_r = 0$$

where \vec{F} contains gravity, the normal force and friction.

* At the point where the bug can barely hold on, the friction force will be on the threshold in balancing the centrifugal and the coriolis forces as shown to the left.



(2)

- $\vec{\omega}$ is \perp to the x-y plane, so

$$|F_{\text{centrif}}| = m\omega^2 r$$

$$|F_{\text{Coriolis}}| = 2m\omega v_r$$

$$\text{And, } |F_{\text{friction}}| = \mu mg$$

- So, we have

$$m\omega^2 r = \mu mg \sin \theta \quad (1)$$

$$2m\omega v_r = \mu mg \cos \theta \quad (2)$$

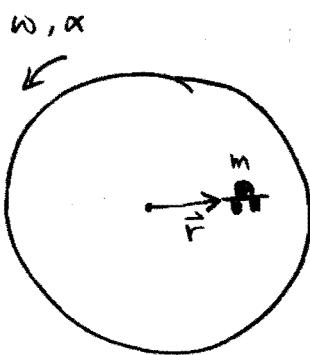
$$(1)^2 + (2)^2 \Rightarrow m^2 \omega^4 r^2 + 4m^2 \omega^2 v_r^2 = \mu^2 m^2 g^2$$

$$r^2 = \frac{\mu^2 g^2 - 4\omega^2 v_r^2}{\omega^4}$$

$$r = \left[\frac{(0.3)^2 (9.81 \text{ m/s}^2)^2 - 4(3.0 \text{ rad/s})^2 (0.5 \times 10^{-2} \text{ m/s})^2}{(3.0 \text{ rad/s})^4} \right]^{1/2}$$

$$\underline{\underline{r = 32.7 \text{ cm}}}$$

4.25



- both $\vec{\omega}$ and $\vec{\alpha}$ are out of the page
 - $\vec{\alpha} = 0.02 \text{ rev/s}^2 = 2\pi(0.02) \text{ rad/s}^2$ (up)
 - constant angular acceleration,
- $$\omega = \omega_0 + \alpha t = \alpha t \text{ (up)} (\omega_0 = 0)$$

- In the rotating frame, we have the following equation of motion:

since $\vec{v}_r = 0$

$$\vec{F}_{\text{eff}} = \vec{F} - 2m(\vec{\omega} \times \vec{v}_r) - m\vec{\omega} \times (\vec{\omega} \times \vec{r}) - m(\vec{\omega} \times \vec{r}) = m\vec{a}_r = 0.$$

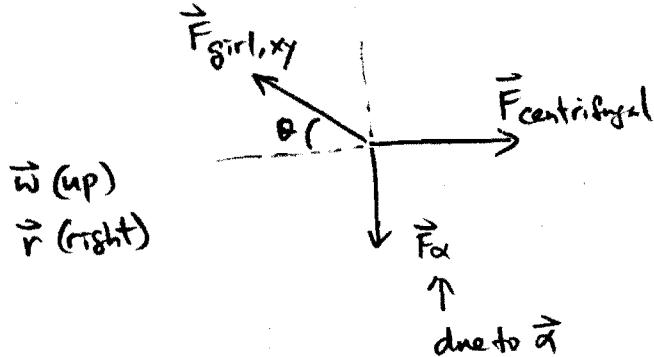
\downarrow
 $\vec{\omega} = \vec{\alpha}$

this includes gravity

$$\oplus \vec{F}_{\text{girl}}$$

- Since $\vec{\omega}, \vec{\alpha}$ are \perp to the x-y plane, the \hat{z} dynamic is trivial: $m\vec{g} = \vec{F}_{\text{girl}, z}$

- On the x-y plane, we have the following free body diagram:



$\vec{F}_{\text{girl}, xy}$ - the projection of \vec{F}_{girl} on the x-y plane.

- This is the relevant one since we don't really care about the trivial $\vec{F}_{\text{girl}, z}$ in the \hat{z} direction.

→ We need to find the magnitude of $\vec{F}_{\text{girl,xy}}$ & θ . (2)

To balance the forces, we have

$$F_{\text{girl}} \cos \theta = F_{\text{cent}} = m \omega^2 r \quad \text{--- (1)}$$

$$F_{\text{girl}} \sin \theta = F_a = m \alpha r \quad \text{--- (2)}$$

$$(1)^2 + (2)^2 \Rightarrow$$

$$\begin{aligned} F_{\text{girl}} &= \left[m^2 \omega^4 r^2 + m^2 \alpha^2 r^2 \right]^{1/2} \\ &= mr (\alpha^2 t^4 + \alpha^2)^{1/2} \quad \omega = \alpha t \\ &= mr \alpha (\alpha^2 t^4 + 1)^{1/2} \\ &= (3.0 \text{ kg})(7.0 \text{ m})(2\pi \cdot 0.02 \text{ rad/s}) \left[(2\pi \cdot 0.02)^2 (6.0)^4 + 1 \right]^{1/2} \\ &= (2.64)(4.63) \text{ N} \\ &= \underline{\underline{12.2 \text{ N}}} \end{aligned}$$

$$\begin{aligned} \frac{(2)}{(1)} \Rightarrow \tan \theta &= \frac{m \alpha r}{m \omega^2 r} = \frac{\alpha}{(\alpha t)^2} = \frac{1}{\alpha t^2} \\ &= \frac{1}{(2\pi)(0.02) \text{ rad/s}^2 \cdot (6.0)^2} = 0.2210 \end{aligned}$$

$$\underline{\underline{\theta = 12.5^\circ \quad (167.5^\circ \text{ cc from } \vec{r})}}$$

