

An Introduction to Statistical Mechanics and ①

the Ising Model

- A branch of physics that deals with systems with a large number of interacting elements.

classic examples: gases
fluids
magnets ...

→ Newton's Eggs.
Hamilton Eggs.

- A direct application of Mechanics to each of its individual parts is impossible due to the extreme size of the system - $N \sim 10^{23}$!

- ★ - However, if one looks at the macroscopic properties of these systems, such as P & T , they are well behaved quantities.
⇒ microscopic properties must "average" out to give predictable macroscopic behavior for the whole system.

(2)

- This "averaging" property of large systems give us the ability to describe the system statistically - "Statistical Mechanics".

Levels of complication

"Microcanonical Ensembles" : isolated systems with no exchange of energy or particles with its environment
(E is conserved here)

- The energy E of the system is fixed by the Hamiltonian function $H(\underline{x})$

where \underline{x} is a high dimensional vector that specifies the configuration of the (state) system.

- ★ - For large systems, there might be a large number of degenerate states for a given value of E .

For the sake of simplicity, suppose that the volume V of the system is its only relevant external parameter. An isolated system of this kind consists then of a given number N of particles in a specified volume V , the constant energy of the system being known to lie in some range between E and $E + \delta E$. Probability statements are then made with reference to an ensemble which consists of many such systems, all consisting of this number N of particles in this volume V , and all with their energy lying in the range between E and $E + \delta E$. The fundamental statistical postulate asserts that in an equilibrium situation the system is equally likely to be found in any one of its accessible states. Thus, if the energy of a system in state r is denoted by E_r , the probability P_r of finding the system in state r is given by

$$P_r = \begin{cases} C & \text{if } E < E_r < E + \delta E \\ 0 & \text{otherwise} \end{cases} \quad (6.1.1)$$

where C is a constant. It can be determined by the normalization condition that $\sum P_r = 1$ when summed over all accessible states in the range between E and $E + \delta E$.

An ensemble representing an isolated system in equilibrium consists then of systems distributed in accordance with (6.1.1). It is sometimes called a "microcanonical" ensemble.

6·2 System in contact with a heat reservoir

We consider the case of a small system A in thermal interaction with a heat reservoir A' . This is the situation already discussed in Sec. 3·6 where $A \ll A'$, i.e., where A has many fewer degrees of freedom than A' . The system A may be any relatively small macroscopic system. (For example, it may be a bottle of wine immersed in a swimming pool, the pool acting as a heat reservoir.) Sometimes it may also be a distinguishable microscopic system which can be clearly identified.* (For example, it may be an atom at some lattice site in a solid, the solid acting as a heat reservoir.) We ask the following question: Under conditions of equilibrium, what is the probability P_r of finding the system A in any one particular microstate r of energy E_r ?

This question is immediately answered by the same reasoning as was used in Sec. 3·3. We again assume weak interaction between A and A' so that their energies are additive. The energy of A is, of course, not fixed. It is only the total energy of the combined system $A^{(0)} = A + A'$ which has a constant value in some range between $E^{(0)}$ and $E^{(0)} + \delta E$. The conservation of energy can then be written as

$$E_r + E' = E^{(0)} \quad (6.2.1)$$

where E' denotes the energy of the reservoir A' . When A has an energy E_r ,

* The qualifying remark is introduced because it may not always be possible to label

the reservoir A' must then have an energy near $E' = E^{(0)} - E_r$. Hence, if A is in the one definite state r , the number of states accessible to the combined system $A^{(0)}$ is just the number of states $\Omega'(E^{(0)} - E_r)$ accessible to A' whose energy lies in a range δE near the value $E' = E^{(0)} - E_r$. But, according to the fundamental statistical postulate, the probability of occurrence in the ensemble of a situation where A is in state r is simply proportional to the number of states accessible to $A^{(0)}$ under these conditions. Hence

$$P_r = C' \Omega'(E^{(0)} - E_r) \quad (6.2.2)$$

where C' is a constant of proportionality independent of r . As usual, it can be determined from the normalization condition for probabilities, i.e.,

$$\sum_r P_r = 1 \quad (6.2.3)$$

where the sum extends over all possible states of A irrespective of energy.

Up to now, our discussion has been completely general. Let us now make use of the fact that A is a very much smaller system than A' . Then $E_r \ll E^{(0)}$ and (6.2.2) can be approximated by expanding the slowly varying logarithm of $\Omega'(E')$ about the value $E' = E^{(0)}$. Thus

$$\ln \Omega'(E^{(0)} - E_r) = \ln \Omega'(E^{(0)}) - \left[\frac{\partial \ln \Omega'}{\partial E'} \right]_0 E_r \cdots \quad (6.2.4)$$

Since A' acts as a heat reservoir, $E_r \ll E^{(0)}$ and higher-order terms in the expansion can be neglected. The derivative

$$\left[\frac{\partial \ln \Omega'}{\partial E'} \right]_0 = \beta \quad (6.2.5)$$

is evaluated at the fixed energy $E' = E^{(0)}$ and is thus a constant independent of the energy E_r of A . By (3.3.10) it is just the constant temperature parameter $\beta = (kT)^{-1}$ characterizing the heat reservoir A' . (Physically, this means that the reservoir A' is so large compared to A that its temperature remains unaffected by whatever small amount of energy it gives to A .) Hence (6.2.4) becomes

$$\ln \Omega'(E^{(0)} - E_r) = \ln \Omega'(E^{(0)}) - \beta E_r$$

or

$$\Omega'(E^{(0)} - E_r) = \Omega'(E^{(0)}) e^{-\beta E_r} \quad (6.2.6)$$

Since $\Omega'(E^{(0)})$ is just a constant independent of r , (6.2.2) becomes then simply

$$P_r = C e^{-\beta E_r} \quad (6.2.7)$$

where C is some constant of proportionality independent of r . Using the normalization condition (6.2.3), C is determined by the relation

$$C^{-1} = \sum_r e^{-\beta E_r}$$

"Canonical ensembles" : systems are in thermal contact with a heat reservoir. (3)

- heat reservoir is assumed to be big so that $T_{\text{reservoir}}$ can be considered fixed

But, energy E can be exchanged between the reservoir and the system.

- At a given temperature T , the system typically has access to a large # of allowed states!

Statistical Description of states

① let define

$R(x \rightarrow x') dt$ to be the probability that the system starting in state x and ending up in state x' after dt .

$R(x \rightarrow x')$ is called the transition rate.

(4)

- $R(\underline{x} \rightarrow \underline{x}')$ is usually assumed to be time independent!

② Now, we can also define a set of weights $w_{\underline{x}}(t)$ that gives the probability in finding the system at state \underline{x} at time t .

③ Then, the time evolution of the system can be describe by the "master equation":

$$\frac{dw_{\underline{x}}}{dt} = \sum_{\underline{x}'} [w_{\underline{x}'}(t) R(\underline{x}' \rightarrow \underline{x}) - w_{\underline{x}}(t) R(\underline{x} \rightarrow \underline{x}')]]$$

- this is just a statement on the conservation of probability.

Since, $w_{\underline{x}}(t)$ is a probability, it must also satisfy:

$$\sum_{\underline{x}} w_{\underline{x}}(t) = 1, \quad \text{for all time.}$$

(5)

★ These two equations formally define $W_x(t)$!

④ With $W_x(t)$ known, then all macroscopic properties can be estimated.

For example, $E(x)$ is the energy for a given state, then

$$\langle E \rangle = \sum_x E(x) W_x(t)$$

sum over all allowed states within the five constraints.

gives the expectation of the energy for the system.

Equilibrium

Consider the situation when the rates of going into & out of x exactly cancel,

then $\frac{dW_x(t)}{dt} = 0$ or $W_x(t)$ is constant in time.

- We would call the system to be in equilibrium.

☆ Since the master equation is of 1st order and $W_x(t) \in [0, 1]$ (probability), all systems governed by this equation must \rightarrow equilibrium at large t . ⑥

☆☆ For systems in equilibrium, $R(x \rightarrow x')$ must take on specific values!

☆☆ For systems in equilibrium, we know how the equilibrium occupation probability

$$P(x) = \lim_{t \rightarrow \infty} W_x(t)$$

should behave. (Gibb 1902, Boltzmann)

\leftarrow Boltzmann's constant

$$\rightarrow P(x) = \frac{1}{Z} e^{-E(x)/KT} \quad \leftarrow \beta = \frac{1}{KT}$$

$$\text{where } Z = \sum_x e^{-E(x)/KT} = \sum_x e^{-\beta E(x)}$$

is called the partition function.

$- Z$ is also the normalization factor.

- $P(x)$ is called the Boltzmann's Distribution

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So, in terms of $P(x)$, macroscopic averages are,

$$\langle Q \rangle = \sum_x P(x) Q(x) = \frac{1}{Z} \sum_x Q(x) e^{-\beta E(x)}$$

In particular, the expectation value of the energy

$\langle E \rangle$ (the internal energy of the system) is

$$U = \frac{1}{Z} \sum_x E(x) e^{-\beta E(x)}$$

Note: This can also be written in terms of Z ,

$$U = - \frac{1}{Z} \frac{\partial Z}{\partial \beta} = - \frac{\partial \log Z}{\partial \beta}$$

check: $\frac{\partial Z}{\partial \beta} = \frac{\partial}{\partial \beta} \left(\sum_x e^{-\beta E(x)} \right)$

$$= \sum_x (-E(x) e^{-\beta E(x)})$$

$$- \frac{1}{Z} \frac{\partial Z}{\partial \beta} = \frac{1}{Z} \sum_x E(x) e^{-\beta E(x)}$$

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And, specific heat is,

$$dT = -\frac{1}{k\beta^2} d\beta$$

$$u = -\frac{\partial \log Z}{\partial \beta}$$

$$C \equiv \frac{\partial U}{\partial T} = -k\beta^2 \frac{\partial U}{\partial \beta} = k\beta^2 \frac{\partial^2 \log Z}{\partial \beta^2}$$

$$\text{recall } \beta = \frac{1}{kT}$$

Then, from thermodynamic relations, we have

$$C = T \frac{\partial S}{\partial T} = -\beta \frac{\partial S}{\partial \beta}$$

$$\begin{aligned} \text{So, } S &= \int -\frac{C}{\beta} d\beta \\ &= -\int k\beta \frac{\partial^2 \log Z}{\partial \beta^2} d\beta \end{aligned}$$

$$u = k\beta \quad dv = \frac{\partial^2 \log Z}{\partial \beta^2} d\beta$$

$$du = k d\beta \quad v = \frac{\partial \log Z}{\partial \beta}$$

$$= -k\beta \frac{\partial \log Z}{\partial \beta} + \int \frac{\partial \log Z}{\partial \beta} k d\beta$$

$$S = -k\beta \frac{\partial \log Z}{\partial \beta} + k \log Z$$

(9)

Then, the free energy is given by

$$F = U - TS = - \frac{\partial \log Z}{\partial \beta} - \frac{1}{k\beta} \left(-k\beta \frac{\partial \log Z}{\partial \beta} + k \log Z \right)$$

$$= -kT \log Z$$

Then, with F , we can calculate all the "conjugate" forces to the system:

$$P = - \frac{\partial F}{\partial V}$$

(P, V)

$$M = \frac{\partial F}{\partial B}$$

(M, B)

★ Z (the partition function) is an important quantity in statistical physics.

In numerical simulations, it is important to be able to calculate Z efficiently!

- All macroscopic thermodynamic quantities can be derived by taking the appropriate derivatives of Z with respect to β , V , B , etc. -

Calculating C more directly:

- Consider the variance of E :

$$\begin{aligned} & \langle (E - \langle E \rangle)^2 \rangle \\ &= \langle E^2 - 2E\langle E \rangle + \langle E \rangle^2 \rangle \\ &= \langle E^2 \rangle - 2\langle E \rangle^2 + \langle E \rangle^2 \\ &= \langle E^2 \rangle - \langle E \rangle^2 \end{aligned}$$

$$\begin{aligned} \text{Now, } \langle E^2 \rangle &= \frac{1}{Z} \sum_{\mathbf{x}} E^2(\mathbf{x}) e^{-\beta E(\mathbf{x})} \\ &= \frac{1}{Z} \frac{\partial^2 Z}{\partial \beta^2} \end{aligned}$$

$$\langle E \rangle^2 = \left(\frac{1}{Z} \frac{\partial Z}{\partial \beta} \right)^2$$

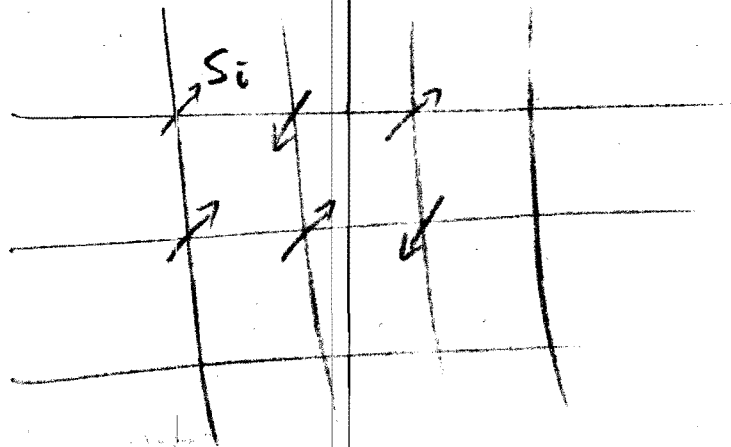
$$\begin{aligned} \text{Note: } \frac{\partial}{\partial \beta} \left(\frac{\partial \log Z}{\partial \beta} \right) &= \frac{\partial}{\partial \beta} \left(\frac{1}{Z} \frac{\partial Z}{\partial \beta} \right) \\ &= \frac{1}{Z} \frac{\partial^2 Z}{\partial \beta^2} - \frac{1}{Z^2} \frac{\partial Z}{\partial \beta} \frac{\partial Z}{\partial \beta} \\ &= \frac{1}{Z} \frac{\partial^2 Z}{\partial \beta^2} - \langle E \rangle^2 \end{aligned}$$

\uparrow \uparrow
 $\frac{C}{k\beta^2}$ $\langle E^2 \rangle$

The Ising Model

- A simple model for a magnet.

Model



- Macroscopic magnetism is resulted from a collection of magnetic dipoles or atomic spins S_i .

- The spins are assumed to be on a fixed lattice (3D physical system / 2D simplified system chain)

Note : 1D & 2D can be solved analytically
(Onsager 1944)
3D is still unresolved analytically.

- S_i can be up (+1) or down (-1)

(other types exist such as

continuous spin models : XY 2D Heisenberg 2D)

- interactions are of the nearest neighbor type.

→ This model provides a picture on how the system phase transition from a paramagnetic state to a ferromagnetic state as $T < T_c$!

paramagnetic → ferromagnetic → }
 Magnetization = 0 in zero field } $\langle M \rangle \neq 0$

The system is characterized by the Hamiltonian

$$H = -J \sum_{\langle ij \rangle} S_i S_j - B \sum_i S_i$$

with all other

it with external B field

for the simulation, take $B=0$.

summing over NN

Then, the partition function is given by

$$Z = \sum_{\{S_i\}} e^{-\beta H}$$

↑
Summing all possible combinations of $\{S_i\}$
with a given energy $E (= H)$.

$$- \langle M \rangle = \left\langle \sum_i S_i \right\rangle_{\text{over ensemble}}$$

mean magnetization

↑
over all spins on one grid

$$- \langle m \rangle = \frac{\langle M \rangle}{N}$$

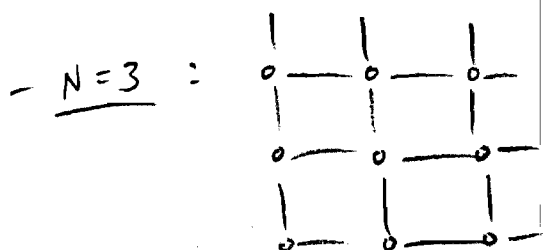
mean magnetization per spin

Counting Bonds on Ising Grid

$$H = -J \sum_{\langle ij \rangle} S_i S_j \quad - \text{sum over all bonds}$$

with periodic boundary

For $N \times N$ grid, $H = -2N^2$ for all up or down arrangement.



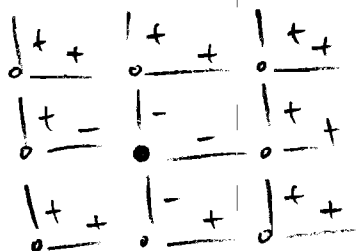
3x3 vertical bonds
3x3 horizontal bonds
2x3x3 total bonds

$$H = -18$$

- With one spin flip at middle

$$E_x' = E_x + 2J S_{ij}^x \sum_{nn} S_{ij}^x$$

$$= -18 + 8 = -10$$



6 actual count = 10

0

4

$$E_x' = -10$$