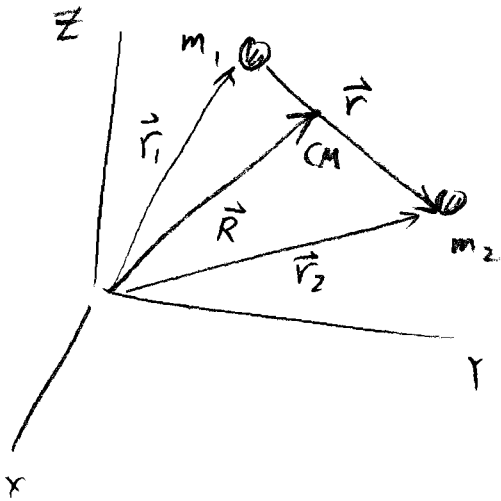


①

Central Force

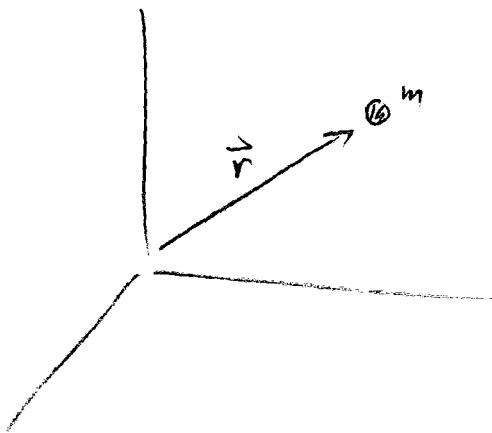


— central force :

interaction force is a function of $\vec{r}$ only!

★ It can be shown that central force motion between (see ①) m_1 & m_2 about their center of mass can always be reduced to an equivalent one body problem of a reduced mass

$m = \frac{m_1 m_2}{m_1 + m_2}$ at a distance $\vec{r}$ from the fixed center of force



①

$$\left\{ \begin{array}{l} \text{mass 1:} \quad m_1 \ddot{\vec{r}}_1 = \vec{F}_{12} \\ \text{mass 2:} \quad m_2 \ddot{\vec{r}}_2 = \vec{F}_{21} \end{array} \right.$$

- Add these two equations:

$$m_1 \ddot{\vec{r}}_1 + m_2 \ddot{\vec{r}}_2 = \vec{F}_{12} + \vec{F}_{21} \quad (1)$$

Note: - the location of the center of mass is given by

$$\vec{R} = \frac{m_1 \vec{r}_1 + m_2 \vec{r}_2}{m_1 + m_2} \quad \& \quad \ddot{\vec{R}} = \frac{m_1 \ddot{\vec{r}}_1 + m_2 \ddot{\vec{r}}_2}{m_1 + m_2}$$

- By Newton's 3rd law, we have $\vec{F}_{12} = -\vec{F}_{21}$
(force along $\vec{r} = \vec{r}_1 - \vec{r}_2$)

Then, Eq (1) becomes:

$$(m_1 + m_2) \ddot{\vec{R}} = 0 \quad \leftarrow \quad \vec{F}_{12} - \vec{F}_{21} \quad \text{gives} \quad \ddot{\vec{R}} = 0$$

$\Rightarrow$ the center of mass is moving at constant velocity.

- Take ref. frame, in which the cm is at rest.

- Now, let consider the relative motion between m_1 & m_2 , (1/1)

$$\begin{aligned}\ddot{\vec{r}} &= \ddot{\vec{r}}_1 - \ddot{\vec{r}}_2 = \frac{\vec{F}_{12}}{m_1} - \frac{\vec{F}_{21}}{m_2} \\ &= \left(\frac{1}{m_1} + \frac{1}{m_2} \right) \vec{F}_{12}, \quad (\vec{F}_{12} = -\vec{F}_{21}) \quad (2)\end{aligned}$$

- With $m = \frac{m_1 m_2}{m_1 + m_2}$ (reduced mass),

$$(2) \Rightarrow \underline{m \ddot{\vec{r}} = \vec{F}_{12}(\vec{r})}$$

- The two body problem can be reduced to an equivalent one body problem with the reduced mass m !

(2)

- In addition, we assume the force is conservative and it depends on $r = |\vec{r}|$ only, i.e.,

$$f(r) = - \frac{\partial U}{\partial r}$$

- Then, we have the following two properties:

① force is conservative, E is conserved.

② $U(r)$ depends on r only, soln will be spherically symmetric $\rightarrow$ $\vec{L}$ is conserved.

$\rightarrow$ if $\vec{L}$ is conserved, $\vec{r}$ must always $\perp$ to $\vec{L}$!
meaning that motions must be in a plane.

- From energy conservation, we have

$$E = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2) + U(r)$$

$$\frac{d}{dt} E = 0 = \frac{d}{dt} \left(\frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2) + U(r) \right)$$

$$\frac{d}{dt} \left(\frac{1}{2} m \dot{r}^2 \right) = - \frac{d}{dt} \left(U + \frac{1}{2} m r^2 \dot{\theta}^2 \right)$$

LHS:

(3)

$$\frac{d}{dt} \left(\frac{1}{2} m \dot{r}^2 \right)$$

$$= \frac{1}{2} m (2 \dot{r}) \ddot{r}$$

$$= m \dot{r} \ddot{r}$$

Note: $\left[\frac{m r^2 \dot{\theta}^2}{2} = \frac{1}{2 m r^2} \overbrace{m^2 r^4 \dot{\theta}^2}^{l^2 - \text{magnitude of angular momentum } \vec{L}} \right]$
 (Note: l is also conserved)

$$= \frac{l^2}{2 m r^2}$$

RHS: $-\frac{d}{dt} \left(U + \frac{1}{2} m r^2 \dot{\theta}^2 \right) = -\frac{d}{dr} (\quad) \frac{dr}{dt}$

$$= - \left[\frac{d}{dr} \left(U(r) + \frac{l^2}{2 m r^2} \right) \right] \dot{r}$$

$$= \left(f(r) - \frac{l^2}{m r^3} \right) \dot{r}$$

Putting them back together, we have,

$$m \dot{r} \ddot{r} = \left(-\frac{l^2}{m r^3} + f(r) \right) \dot{r}$$

$$\underline{m \ddot{r} = -\frac{l^2}{m r^3} + f(r)} \quad \underline{\text{equation of motion}}$$

(We will integrate this later. Now, we want to see what we can say by analyzing the $U'(r)$ term.)